

# Gauge Symmetries for Spin-2 Theories

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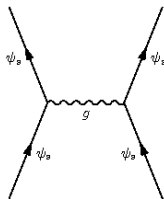
Based partly on [hep-th/0606019](https://arxiv.org/abs/hep-th/0606019)  
*E. Álvarez (UAM), D.B, J.Garriga (UB) & E. Verdaguer (UB)*

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- 1 Massless Spin-2
  - Massless Spin-2 and Gravitation
- 2 Transverse Gauge Symmetry and Spin-2
  - Well-behaved Lagrangians
  - Pure Spin-2 Lagrangians
- 3 Non-linear Extension

# Massless Spin-2 and Gravitation

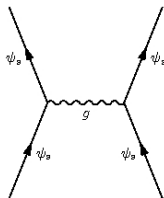
We expect *gravity* to be mediated by a **massless** field  $g$  of a given **spin**.



- Fermions are ruled out ( $\Delta J = \pm 1/2$  or wrong radial dependence)
- Spin 0 boson does not deviate light (couples to the trace)
- Spin  $2n - 1$ ,  $\forall n \in \mathbb{Z}$  alike charges repel.
- **Linear Spin 2** wrong Mercury perihelion (corrected by non-linearities).

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# TDiff Gauge Symmetry and Massless Spin-2

- Massless particle of spin-2: **2 polarizations**.
- To describe them with a tensor  $h_{\mu\nu} = 2 \oplus 1 \oplus 0 \oplus 0$ , a **gauge symmetry** is required.
- The minimal is **TRANSVERSE** (linear) **Diff**.

$$h_{\mu\nu} \sim h_{\mu\nu} + \partial_{(\mu}\xi_{\nu)}, \quad \partial^\mu \xi_\mu = 0$$

van der Bij, van Dam & Ng, 81

They leave  $h = \eta^{\mu\nu} h_{\mu\nu}$  invariant.

# TDiff Lagrangians from Consistency

The most general covariant **massless** lagrangian for  $h_{\mu\nu}$  is

$$\begin{aligned}\mathcal{L} &= \frac{1}{4}\partial_\mu h^{\nu\rho}\partial^\mu h_{\nu\rho} - \frac{\beta}{2}\partial_\mu h^{\mu\rho}\partial_\nu h^\nu_\rho \\ &+ \frac{a}{2}\partial^\mu h\partial^\rho h_{\mu\rho} - \frac{b}{4}\partial_\mu h\partial^\mu h\end{aligned}$$

D.o.f

$$h_{00} = 2A$$

$$h_{0i} = \partial_i B + V_i$$

$$h_{ij} = -2\psi\delta_{ij} + 2\partial_i\partial_j E + 2\partial_{(i}F_{j)} + t_{ij}$$

with  $\partial^i F_i = \partial^i V_i = \partial^i t_{ij} = t_j^j = 0$ .

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# Ghost and Tachyon Free TDiff Lagrangians

$$\begin{aligned}\mathcal{L}_m &= \frac{1}{4} \partial_\mu h^{\nu\rho} \partial^\mu h_{\nu\rho} - \frac{1}{2} \partial_\mu h^{\mu\rho} \partial_\nu h^\nu_\rho \\ &+ \frac{a}{2} \partial^\mu h \partial^\rho h_{\mu\rho} - \frac{b}{4} \partial_\mu h \partial^\mu h - \frac{1}{2} m^2 h^2\end{aligned}$$

$\psi$  : Massive Spin-0 of mass  $m$ .

$t_{ij}$  : Massless Spin-2.

For  $n \neq 2$ ,

$$b \leq \frac{1 - 2a + (n-1)a^2}{(n-2)}, \quad m^2 > 0.$$

- Phenomenology: standard scalar-tensor ( $m \gtrsim (30\mu m)^{-1}$ ) Will, 05
- Pure spin-2

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van Nieuwenhuizen, 73

# Pure Spin-2 Lagrangians

**Almost** all come from usual **Fierz-Pauli** ( $a = b = 1$ ) after field redef.  
**One** pathological case:

Two **INEQUIVALENT** lagrangians which only propagate massless spin-2 (and enlarged **Gauge Symmetry**)!!:

- Fierz-Pauli (GR):  $a = b = 1$  (and field redefinitions).

**G. Symmetry:** Diff

$$\delta h_{\mu\nu} = 2\partial_{(\mu}\xi_{\nu)}.$$

Only possibility for **massive** spin-2.

van Nieuwenhuizen, 73

- WTDiff:  $a = \frac{2}{n}$ ,  $b = \frac{n+2}{n^2}$ .

**G. Symmetry:** Weyl & TDiff

$$\delta h_{\mu\nu} = 2\partial_{(\mu}\xi_{\nu)}, \quad \partial^\mu \xi_\mu = 0, \quad \delta h_{\mu\nu} = 2\eta_{\mu\nu}\phi.$$

# TDiff Non-Linear Extension

From the linear symmetries:

- $f(\eta_{\mu\nu}, g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}) = h + O(h^2)$  as an **invariant**.

$$S[g_{\mu\nu}, \psi, \eta_{\mu\nu}] = \int d^n x L(f, g_{\mu\nu}, \psi)$$

# TDiff Non-Linear Extension

From the linear symmetries:

- Most natural possibility:  $f = g \equiv \det g_{\mu\nu}$

$$S[g_{\mu\nu}, \psi] = \int d^n x (-\Phi(g, \psi) R(\hat{g}_{\mu\nu}) + L[g, \psi, \hat{g}_{\mu\nu}])$$

$$\hat{g}_{\mu\nu} = |g|^{-1/n} g_{\mu\nu}.$$

- Invariant under Diff. of  $J = \det \frac{\partial x}{\partial y} = 1$ . (TDiff)
- Constraint by the linear analysis.

## E.o.m.: Scalar-Tensor.

Einstein's frame:  $\bar{g}_{\mu\nu} = \Phi \hat{g}_{\mu\nu}$

Buchmuller & Dragon , 88

$$G_{\mu\nu}(\bar{g}_{\mu\nu}) = \kappa \bar{T}_{\mu\nu}(g, \psi) + \Lambda \bar{g}_{\mu\nu}.$$

- $g$  as a scalar.
- $\Lambda$  is an integration constant.

# Non-Linear Extension: Pure GR

- Weyl Symmetry ( $g_{\mu\nu} \mapsto \Omega^2 g_{\mu\nu}$ ).  $\delta_W \hat{g}_{\mu\nu} = \delta_W g^{-1/n} g_{\mu\nu} = 0$ .

$$S[g_{\mu\nu}, \psi] = \int d^n x (-\hat{g}^{\mu\nu} R_{\mu\nu}(\hat{g}_{\mu\nu}) + L_W[g, \psi, \hat{g}_{\mu\nu}])$$

~ Unruh, 89

E.o.m.

$$R_{\mu\nu} - \frac{1}{n} g_{\mu\nu} R = f_{\mu\nu}(g)$$

In the gauge  $|g| = 1$ ,  $f_{\mu\nu}(g) = 0$  thus using **geometrical Bianchi**:

$$2\nabla^\mu R_{\mu\nu} = \nabla_\nu R, \implies \nabla_\nu R = 0$$

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = \Lambda g_{\mu\nu}$$

- $S = \int (\sqrt{-g} R + f(g)) d^n x$  (SSB?, Anomalies?)
- Other  $\eta_{\mu\nu}$  dependent possibilities.

# Non-Linear Extension (Gupta's Program)

Hope: Finding a **non-linear** extension through **self-interaction**.

$$\text{E.o.m. } \mathcal{D}(h)_{\mu\nu} = \frac{\delta \mathcal{L}_{(m=0)}}{\delta h^{\mu\nu}} = 0$$

$$\text{TDiff: } \partial^\mu \mathcal{D}(h)_{\mu\nu} = \partial_\nu \Phi$$

$$\text{WTDiff: } \partial^\mu \mathcal{D}(h)_{\mu\nu} = \partial_\nu \Phi, \quad \eta^{\mu\nu} \mathcal{D}(h)_{\mu\nu} = 0.$$

The natural equation to consider is  $T_{\mu\nu}(h) + \beta T(h)\eta_{\mu\nu} = \mathcal{D}(h)_{\mu\nu}$ .

But:

- Wald's (also Damour, Henneaux et al.) extension not clear.
- Deser's method OK but not what we expect.
- It is not clear how to extend **uniquely** non-linearly  $\partial_\mu \xi^\mu = 0$ .
- $h$  is the **linearization** of  $\det(g)$ , but **also** of  $\eta^{\mu\nu} h_{\mu\nu}$ .

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# Conclusions

- The **minimal gauge symmetry** for **massless spin-2** is **TDiff**:

$$\delta h_{\mu\nu} = 2\partial_{(\mu}\xi_{\nu)}, \quad \partial^\mu \xi_\mu = 0$$

- Massless Spin-2 admits naturally a (well-behaved) **scalar** partner, except
  - Fierz-Pauli (**Diff** invariant).
  - WTDiff (**Weyl & TDiff** invariant).
- We found (not unique) **non-linear completions** equivalent to **Scalar-Tensor**. Restricting to GR: **Diff** and **WTDiff** have equivalent **e.o.m.** except for the origin of  $\Lambda$ .
- **Quantization** (BRST) is possible and the **anomalies** are similar to the usual gravitational anomalies. Still maybe **quantum** differences.